Crystallization Conditions and Mineral Chemistry in the East of Tafresh, Central Iran, with Insights into Magmatic Processes



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Abstract: The Tafresh granitoids are located at the central part of the Urumieh-Dokhtar Magmatic Arc (UDMA) in Iran. These rocks, mainly consisting of diorite and granodiorite, were emplaced during the Early Miocene. They are composed of varying proportions of plagioclase + K-feldspar + hornblende \pm quartz \pm biotite. Discrimination diagrams and chemical indices of amphibole phases reveal a calc-alkaline affinity and fall clearly in the crust-mantle mixed source field. The estimated pressure, derived from Al in amphibole barometry, is approximately 3 Kb. The granitoids are I-type, metaluminous and belong to the calc-alkaline series. They are all enriched in light rare earth elements and large ion lithophile elements, depleted in high field strength elements and display geochemical features typical of subduction-related calc-alkaline arc magmas. Most crystal size distribution (CSD) line patterns from the granitoids show a non-straight trend which points to the effect of physical processes during petrogenesis .The presence of numerous mafic enclaves, sieve texture and oscillatory zoning along with the CSD results show that magma mixing in the magma chamber had an important role in the petrogenesis of Tafresh granitoids. Moreover, the CSD analysis suggests that the plagioclase crystals were crystallized in a time span of less than 1000 years, which is indicative of shallow depth magma crystallization.

Key words: crust-mantle mixed source, I-type, subduction, crystal size distribution, magma mixing, Urumieh-Dokhtar Magmatic Arc

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1 Introduction

The behavior of magma in the chamber is controlled by different factors such as: temperature, pressure, chemistry, oxygen fugacity, etc., the variations of any of which will affect the texture of the resultant rock. Therefore, the texture of an igneous rock is a remarkable recorder of igneous processes (Van der Zwan et al., 2013; Viccaro et al., 2006).

As igneous rocks crystallize, there occurs a complex interplay of crystal nucleation and growth. Whether in relatively quickly cooled volcanic rocks and shallow intrusions, or in slowly cooled plutonic rocks, these changes are recorded in their textures. A key to understanding the formation of igneous rocks is to read and interpret their crystal chemical compositions and textures (Renzulli and Santi, 1997; Jerram and Cheadle, 2000; Kuşcu and Floyd, 2001). Quantitative analyses of textures (i.e., crystal size, form, orientation, and association) are a basis for understanding the physical characteristics of magma crystallization (Higgins, 2006; Mock et al., 2003; Marsh, 1988; Jerram and Higgins, 2007). Analysis of CSD, introduced by Randolph and Larson (1988), is a quantitative petrographic technique (e.g., Higgins, 1988; 2006; 2011) that consists of plotting the natural logarithm of the population density (n) vs. crystal length (L) (Castro et al., 2004). In general, as crystal size increases, population density decreases. A typical CSD plot is a linear, or gently curved array with a negative slope (Cashman and Marsh, 1988; Marsh, 1988; 1998; Higgins, 1996; 2000; 2002). Analytical expressions of crystal size, population density and crystal volume can be derived from the regression trend of the CSD array, as discussed by many authors (e.g., Cashman and Marsh, 1988; Cashman, 1990; Higgins, 1994; 2000; 2006; Mock and Jerram, 2005; Jerram and Higgins, 2007;Pourkhorsandi et al., 2015).

In this study, the CSD technique is used to provide insight into the relative roles of residence time, textural coarsening, magma mixing, and renewed magma injection during cooling of Miocene granitoids in the Tafresh granitoids of the Urumieh-Dokhtar Magmatic Arc (UDMA), central Iran.

The principal objective of this study is to explore the mineralogical evidence that provides insight into potential magma mixing/mingling processes in the Tafresh granitoids. The Tafresh granitoids, of Miocene age, immediately followed widespread Eocene magmatic activity. Tafresh diorite (20.1 mya) and granodiorite (19.4 mya) (Mirnejad et al., 2018) are I-type granitoids similar

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to several other intrusive rocks of the UDMA. To get insights to the preceding questions, we tried to quantify the textural information and mineral chemistry recorded in these granitoids with an emphasis on crystal size distributions of plagioclase, which is ideal for this purpose because of the abundance and wide compositional variation of plagioclase.

2 Geological Background and Petrography

Tafresh granitoids are situated in the UDMA 24 km to the east of the city of Tafresh (Fig. 1). Igneous emplacement was associated with the northward subduction of the Neo-Tethys oceanic crust beneath the Iranian continental plate (Stöklin, 1968; Berberian et al., 1982; Alavi, 1994). Deposition of felsic to intermediate volcanics (Agard et al, 2011) was followed by the main plutonic activity (emplacement of batholiths and porphyritic stocks) during the Eocene, 55–36 my ago (Verdel et al., 2011). The volcanic rocks near Tafresh consist of rhyolite, dacite, andesite and basalt, which are locally interbedded with sediments. These sediments contain shallow marine fossils, and display textures suggestive of submarine eruption (Hajian, 2001).

Tafresh granitoids occupy a 11×5 km area, trending generally E-W, extending from Sarbadan to Ghahan (Fig. 2). Intrusive and sub-volcanic bodies of diorite to

granodiorite, with blocky and massive structures, are in contact with younger sedimentary units and Eocene volcanic rocks (Fig. 2).

The Sarbadan 5×4.5 km body, mostly granodiorite, has a medium-grained and granular texture that consists of plagioclase (45–50%), K-feldspar (30–35%), quartz (10– 20%), biotite (5–10%) and hornblende (5–10%) (Fig. 3a, b). Accessory minerals are opaque minerals (2–3%), pyroxene, zircon and titanite (1–2%). Plagioclase forms euhedral to subhedral and tabular crystals (0.2 to 2.7 mm) with optical zonation (Fig. 3a) and sieve texture (Fig. 3b). Most of the quartz occurs as small interstitial grains (0.1 to 0.64 mm). Orthoclase ranges from 0.3 to 2.2 mm in size (Fig. 3a). Idiomorphic amphibole grains are scattered within a feldspar and quartz matrix. Amphibole crystals, from 2.5 to 0.1 mm, show poikilitic textures containing plagioclase \pm opaque minerals (Fig. 3c).

The Ghahan 4×5 km igneous body consists of diorite and lesser quantities of quartz diorite, having a subgranular to granular texture (Fig. 3d). The minerals present are mainly amphibole, plagioclase, \pm alkali feldspar, \pm quartz (Fig. 3d). Plagioclase (65–85%) is mostly idiomorphic to hypidiomorphic, and elongated (3.12–0.13 mm). Some plagioclase crystals exhibit sieve texture (Fig. 3e). Orthoclase (10–20%) occurs as euhedral to anhedral crystals (1.2-0.2 mm), partly altered to sericite. The groundmass is composed of intergranular fine-grained (0.6



Fig. 1. Simplified geological map of Iran (Aghanabati, 2004).



Fig. 2. Simplified geological map of the Tafresh area (Hajian, 2001). Ghahan and Sarbadan stocks are indicated by a dark box.

-0.1 mm) quartz (5–15%) (Fig. 3a, b). Amphibole (5–15%) occurs as euhedral to subhedral phenocrysts (3.6–0.2 mm) partially altered to chlorite and with a poikilitic texture (Fig. 3d). Sporadic opaque minerals (2–5%) occur as small grains (0.01-0.4 mm) either in the groundmass, or as inclusions in larger crystals.

Independent and agglomerated globular or elliptical microgranular mafic enclaves (MMEs) are locally abundant and occur throughout the Tafresh granitoids. The Tafresh granitoids contain many independent and agglomerated globular or elliptical microgranular mafic enclaves (MMEs) that are dark colored in hand-specimens. These enclaves display sharp contacts with the host MMEs in the Tafresh granitoids showing common magmatic textures and are ellipsoidal, with sizes ranging from 0.5 to 5 cm, and sharp contacts against the host granitoid (Fig. 4). Petrographically, the MMEs are finegrained and display distinctive mineral assemblages that are fine grained, and mainly consist of plagioclase and amphibole (Fig. 4). The groundmass consists of plagioclase (65–75%), hornblende (10–15%), orthoclase (5-10%) and subordinate amounts of opaque minerals (2-5%). Euhedral crystals of plagioclase enclose small grains. rounded hornblende Typically, subhedral hornblende ranges from 0.3 to 0.7 mm. The accessory minerals, including epidote, apatite laths, and titanite, are observable in most granitoid samples .The presence of acicular apatite and plagioclase crystals with oscillatory zoning, and resorption surfaces from mixing textures can be interpreted as other non-equilibrium textures in the studied rocks (e.g., Barbarin, 1990; Hibbard, 1991, 1995).

3 Analytical Methods

Electron microprobe analysis of amphibole and

plagioclase grains was done at the Iranian Mineral Processing Research Center using a CAMECA SX 100 equipped with a wavelength-dispersive spectrometer (WDS). The accelerating voltage was 15 kV, with a beam current of 20 nA and a 0-2 µm focused electron beam. The data were corrected using a modified ZAF correction procedure (Fialin, 1988). The precision of the method for major elements was <0.5% of the measured concentration. Tables 1-3 provide representative amphibole and plagioclase analyses. The beam sizes used for phenocrysts and matrix were 3 µm and 1 µm, respectively. The natural and synthetic standards used were minerals of known compositions included albite (Na), periclase (Mg), corundum (Al), wollastonite (Ca, Si), orthoclase (K), pyrophanite (Mn, Ti) and magnetite (Fe). Counting times were 10-30 s for peaks and 10 s for background. Analytical accuracy varies between 1 and 5 % depending on the abundance of the element.

Following petrographic studies, 6 fresh samples were selected for geochemical analysis by inductively coupled plasma optical emission spectroscopy (ICP-OES) at the Department of Geology and Environmental Earth Sciences, Miami University, USA. During preparation, 50 mg of a sample powder was fused with 75 mg LiBO₂ and dissolved in 125 ml of 0.3 N HNO₃. Repeated analyses established a precision of $\pm 1\%$ for major elements, and $\pm 5\%$ for most trace elements (Table 4).

Eight thin sections were selected for quantitative optical microscopy and CSD studies. Mosaic images were obtained at the CEREGE Institute, Aix-Marseille University (France) for identification of the typology and estimation of mineral abundance and size. Measured parameters included length, width, area and crystal orientation. The mosaic images were a basis on which to draw precise images of plagioclase crystals in preparation



Fig. 3. (a) Granular granitoid rock in Ghahan region. Plagioclase crystals are associated with alkali feldspar, quartz, amphibole, and opaque minerals; (b) euhedral plagioclase with sieve texture; (c, d) Poikilitic amphibole (plagioclase inclusions); (e) plagioclase phenocrysts with oscillatory zoning and sieve texture; (f) fine grain microdiorite enclaves.



Fig. 4. (a, b) Fine-grained microdiorite enclaves in diorite.

for the measurement of crystal length and width using ImageJ and JMicroVision 1.2.7 software. The procedure was monitored by continual comparison with optical images and direct microscopic investigation. Finally, the 3 -D characteristics of the crystals were compared with the crystal size distribution diagram, the frequency of the width-to-length ratio of the crystals, and these were compared to the diagrams from the CSDSlice5 Spreadsheet (Morgan and Jerram 2006).

The obtained data were added to a CSD Corrections 1.4.0 software environment to draw CSD diagrams. The input of CSD Corrections requires values on intersection lengths of the crystals, area of the measured thin section, crystal roundness, rock fabric and crystal habit.

As proposed by Marsh (1988), the average value for crystal residence time is calculated by employing the equation:

$n=n_0\exp(-L/Gt)$

where n=dN/dL; n refers to the total number of crystals smaller than L and L is the crystal size in each span. G is the average rate of linear growth, t is the average time for the presence of the crystal in the system and n_0 refers to

| Table | 1 Re | presen | ntative | electro | n micr | oprobe | analys | es of an | phibole | s from | the Gh | ahan an | ıd Sarba | dan gran | itoids(v | vt%) | | | | | | | |
|----------|-------|--------|--------------------------------|------------------|--------|--------|--------|----------|---------|--------|--------|----------|----------|-------------------|----------|-----------------|----------------|-----------|------|-------|------|----------------|---------|
| Sam | le | SiO₂ ≠ | Al ₂ O ₃ | TiO ₂ | FeO | MnO | MgO | CaO | Na_2O | K_2O | NiO | P_2O_5 | Total | No. of oxygens | TSi | Mg /(Mg+Fe2) | (Ca+Na) (B) | Na (B) | Al t | Al vi | Aliv | Fe/ (Fe+Mg) | log fO2 |
| | Gh1 | 49.3 | 6.2 | 1.5 | 11.9 | 0.4 | 14.9 | 11.5 | 1.3 | 0.6 | 0.0 | 0.0 | 97.5 | 23.0 | 7.1 | 0.8 | 2.0 | 0.2 | 1.1 | 0.2 | 0.9 | 0.3 | -12.8 |
| - | Gh2 | 49.3 | 6.4 | 1.4 | 11.9 | 0.4 | 15.0 | 11.5 | 1.1 | 0.7 | 0.0 | 0.1 | 97.8 | 23.0 | 7.1 | 0.8 | 2.0 | 0.2 | 1.1 | 0.2 | 0.9 | 0.3 | -12.6 |
| - | Gh3 | 49.9 | 6.3 | 1.4 | 12.0 | 0.5 | 14.8 | 11.4 | 1.7 | 0.8 | 0.0 | 0.0 | 98.6 | 23.0 | 7.1 | 0.7 | 2.0 | 0.3 | 1.1 | 0.2 | 0.9 | 0.3 | -12.9 |
| - | Gh4 | 49.3 | 5.6 | 1.1 | 13.7 | 0.6 | 14.3 | 11.5 | 1.2 | 0.6 | 0.0 | 0.0 | 97.9 | 23.0 | 7.1 | 0.7 | 2.0 | 0.2 | 1.0 | 0.1 | 0.9 | 0.3 | -13.1 |
| - | Gh5 | 48.6 | 5.7 | 1.5 | 13.6 | 0.5 | 13.5 | 12.5 | 1.2 | 0.6 | 0.0 | 0.0 | 97.8 | 23.0 | 7.1 | 0.6 | 2.0 | 0.0 | 1.0 | 0.1 | 0.9 | 0.4 | -13.2 |
| - | Gh6 | 48.4 | 6.0 | 1.2 | 13.6 | 0.5 | 14.2 | 12.4 | 1.3 | 0.6 | 0.0 | 0.0 | 98.2 | 23.0 | 7.0 | 0.7 | 2.0 | 0.1 | 1.0 | 0.1 | 1.0 | 0.3 | -12.7 |
| - | Gh7 | 49.5 | 5.8 | 1.3 | 14.5 | 0.6 | 13.4 | 12.7 | 0.0 | 0.6 | 0.0 | 0.0 | 98.3 | 23.0 | 7.2 | 0.7 | 2.0 | 0.0 | 1.0 | 0.1 | 0.8 | 0.4 | -13.4 |
| - | Gh8 | 50.7 | 5.3 | 0.9 | 11.7 | 0.6 | 16.6 | 11.6 | 1.1 | 0.4 | 0.0 | 0.0 | 98.9 | 23.0 | 7.1 | 0.9 | 2.0 | 0.3 | 0.9 | 0.0 | 0.9 | 0.3 | -12.3 |
| - | Gh9 | 51.5 | 4.6 | 0.9 | 11.1 | 0.6 | 16.8 | 11.6 | 0.9 | 0.3 | 0.0 | 0.1 | 98.4 | 23.0 | 7.2 | 0.9 | 2.0 | 0.2 | 0.8 | 0.0 | 0.8 | 0.3 | -12.8 |
| <u> </u> | Gh10 | 52.3 | 4.6 | 0.9 | 11.2 | 0.6 | 17.1 | 11.9 | 0.9 | 0.4 | 0.0 | 0.0 | 100.0 | 23.0 | 7.2 | 0.9 | 2.0 | 0.2 | 0.8 | 0.0 | 0.8 | 0.3 | -12.8 |
| | Jh 11 | 47.6 | 7.5 | 1.6 | 12.1 | 0.6 | 16.3 | 11.2 | 1.8 | 0.3 | 0.0 | 0.0 | 99.1 | 23.0 | 6.7 | 0.9 | 2.0 | 0.3 | 1.2 | 0.0 | 1.2 | 0.3 | -10.6 |
| <u> </u> | 3h12 | 48.5 | 7.0 | 1.2 | 12.0 | 0.5 | 17.5 | 10.6 | 1.9 | 0.3 | 0.0 | 0.0 | 99.5 | 23.0 | 6.7 | 1.0 | 1.9 | 0.4 | 1.1 | 0.0 | 1.1 | 0.3 | * |
| u | 3h13 | 50.8 | 4.5 | 0.9 | 10.7 | 0.7 | 19.3 | 10.9 | 1.0 | 0.4 | 0.0 | 0.0 | 99.2 | 23.0 | 7.0 | 1.0 | 1.9 | 0.3 | 0.7 | 0.0 | 0.7 | 0.2 | * |
| eye | 3h14 | 49.3 | 5.7 | 1.2 | 12.1 | 0.7 | 16.5 | 11.2 | 1.3 | 0.5 | 0.0 | 0.0 | 98.5 | 23.0 | 7.0 | 0.9 | 2.0 | 0.3 | 0.9 | 0.0 | 0.9 | 0.3 | -11.2 |
| <u>у</u> | 3h15 | 51.1 | 4.8 | 1.0 | 10.8 | 0.7 | 18.4 | 11.1 | 1.1 | 0.4 | 0.0 | 0.0 | 99.4 | 23.0 | 7.0 | 1.0 | 1.9 | 0.3 | 0.8 | 0.0 | 0.8 | 0.2 | * |
| | Gh16 | 51.3 | 4.7 | 0.9 | 11.2 | 0.7 | 19.2 | 10.9 | 1.0 | 0.4 | 0.0 | 0.0 | 100.2 | 23.0 | 7.0 | 1.0 | 1.9 | 0.3 | 0.8 | 0.0 | 0.8 | 0.2 | * |
| <u> </u> | 3h17 | 51.4 | 4.5 | 1.0 | 11.3 | 0.8 | 17.7 | 11.3 | 1.0 | 0.4 | 0.0 | 0.0 | 99.2 | 23.0 | 7.1 | 1.0 | 1.9 | 0.3 | 0.7 | 0.0 | 0.7 | 0.3 | * |
| <u> </u> | 3h18 | 49.7 | 6.1 | 1.2 | 11.7 | 0.7 | 17.1 | 11.1 | 1.3 | 0.2 | 0.0 | 0.0 | 99.1 | 23.0 | 6.9 | 1.0 | 2.0 | 0.3 | 1.0 | 0.0 | 1.0 | 0.3 | -10.9 |
| <u> </u> | 3h19 | 51.2 | 4.9 | 0.9 | 11.0 | 0.6 | 18.5 | 10.9 | 1.1 | 0.4 | 0.0 | 0.0 | 99.5 | 23.0 | 7.0 | 1.0 | 1.9 | 0.3 | 0.8 | 0.0 | 0.8 | 0.3 | * |
| <u> </u> | 3h20 | 52.0 | 6.1 | 1.5 | 11.7 | 0.5 | 15.4 | 11.3 | 1.5 | 0.4 | 0.0 | 0.0 | 100.4 | 23.0 | 7.2 | 0.8 | 2.0 | 0.3 | 1.0 | 0.2 | 0.8 | 0.3 | * |
| <u> </u> | 3h21 | 51.4 | 6.1 | 1.7 | 13.4 | 0.5 | 14.2 | 11.4 | 1.5 | 0.6 | 0.0 | 0.1 | 100.7 | 23.0 | 7.2 | 0.7 | 2.0 | 0.3 | 1.0 | 0.2 | 0.8 | 0.3 | * |
| <u> </u> | Jh22 | 50.1 | 5.7 | 1.0 | 11.9 | 0.7 | 15.7 | 11.2 | 4.1 | 0.4 | 0.0 | 0.0 | 100.7 | 23.0 | 7.1 | 0.7 | 2.0 | 0.3 | 1.0 | 0.1 | 0.9 | 0.3 | -12.6 |
| <u> </u> | 3h23 | 49.2 | 5.9 | 1.0 | 12.0 | 0.7 | 15.4 | 11.3 | 1.6 | 0.5 | 0.0 | 0.0 | 97.5 | 23.0 | 7.1 | 0.8 | 2.0 | 0.3 | 1.0 | 0.1 | 0.9 | 0.3 | -12.5 |
| <u> </u> | 3h24 | 49.8 | 5.5 | 0.9 | 11.8 | 0.8 | 15.6 | 10.8 | 1.7 | 0.3 | 0.0 | 0.0 | 97.2 | 23.0 | 7.1 | 0.8 | 2.0 | 0.3 | 0.9 | 0.1 | 0.9 | 0.3 | -12.8 |
| <u> </u> | 3h25 | 49.4 | 5.7 | 1.1 | 11.6 | 0.5 | 15.5 | 11.4 | 1.4 | 0.5 | 0.0 | 0.0 | 97.1 | 23.0 | 7.1 | 0.8 | 2.0 | 0.2 | 1.0 | 0.1 | 0.9 | 0.3 | -12.6 |
| <u> </u> | 3h26 | 48.4 | 9.6 | 1.2 | 11.9 | 0.6 | 15.2 | 10.8 | 1.5 | 0.4 | 0.0 | 0.0 | 96.5 | 23.0 | 7.0 | 0.8 | 2.0 | 0.3 | 1.1 | 0.1 | 1.0 | 0.3 | -12.3 |
| | 3h27 | 49.5 | 5.2 | 1.0 | 12.3 | 0.7 | 15.6 | 11.3 | 3.3 | 0.4 | 0.0 | 0.0 | 99.4 | 23.0 | 7.1 | 0.7 | 2.0 | 0.3 | 0.9 | 0.0 | 0.9 | 0.3 | -12.6 |
| | Sal | 49.4 | 5.4 | 1.3 | 14.2 | 0.6 | 14.1 | 11.7 | 1.2 | 0.5 | 0.0 | 0.0 | 98.5 | 23.0 | 7.1 | 0.7 | 2.0 | 0.2 | 0.9 | 0.0 | 0.9 | 0.4 | -13.3 |
| | Sa2 | 48.8 | 5.4 | 1.4 | 13.7 | 0.6 | 14.5 | 11.9 | 1.2 | 0.6 | 0.0 | 0.1 | 98.1 | 23.0 | 7.1 | 0.7 | 2.0 | 0.2 | 0.9 | 0.0 | 0.9 | 0.3 | -12.8 |
| | Sa3 | 48.9 | 6.3 | 1.6 | 13.3 | 0.5 | 14.0 | 11.5 | 1.3 | 0.6 | 0.0 | 0.0 | 98.0 | 23.0 | 7.1 | 0.7 | 2.0 | 0.2 | 1.1 | 0.1 | 0.9 | 0.3 | -13.0 |
| uel | Sa4 | 46.8 | 6.0 | 1.5 | 13.3 | 0.4 | 15.6 | 12.4 | 1.4 | 0.5 | 0.0 | 0.1 | 98.0 | 23.0 | 6.8 | 0.8 | 2.0 | 0.1 | 1.0 | 0.0 | 1.0 | 0.3 | -10.1 |
| psdr | Sa5 | 44.6 | 7.4 | 1.0 | 16.4 | 0.6 | 11.1 | 12.0 | 1.2 | 0.9 | 0.0 | 0.0 | 95.2 | 23.0 | 6.8 | 0.6 | 2.0 | 0.0 | 1.3 | 0.2 | 1.2 | 0.5 | * |
| вS | Sa6 | 49.8 | 5.4 | 1.3 | 13.8 | 0.6 | 14.3 | 12.2 | 1.1 | 0.5 | 0.0 | 0.0 | 99.1 | 23.0 | 7.1 | 0.7 | 2.0 | 0.1 | 0.9 | 0.1 | 0.9 | 0.4 | -13.2 |
| | Sa7 | 49.3 | 5.4 | 1.6 | 13.7 | 0.5 | 14.7 | 12.3 | 1.2 | 0.5 | 0.0 | 0.0 | 99.2 | 23.0 | 7.1 | 0.7 | 2.0 | 0.1 | 0.9 | 0.0 | 0.9 | 0.3 | -12.8 |
| | Sa8 | 49.6 | 5.6 | 1.5 | 12.8 | 0.5 | 14.6 | 11.4 | 1.4 | 0.4 | 0.0 | 0.0 | 97.7 | 23.0 | 7.2 | 0.7 | 2.0 | 0.2 | 1.0 | 0.1 | 0.8 | 0.3 | -13.2 |
| | Sa9 | 49.8 | 5.1 | 0.9 | 12.8 | 0.7 | 15.0 | 11.5 | 0.8 | 0.4 | 0.0 | 0.1 | 97.0 | 23.0 | 7.2 | 0.8 | 2.0 | 0.2 | 0.9 | 0.0 | 0.8 | 0.3 | -13.0 |

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|----------|---------|------------------|------------------|-----------|-------|----------|-----------|---------|-------------------|------------------|-------------------|----------|----------------------|------|-------|-------|
| Sample | No. | SiO ₂ | TiO ₂ | Al_2O_3 | FeO | MnO | MgO | CaO | Na ₂ O | K ₂ O | NiO | P_2O_5 | Total | Or | Ab | An |
| Gh | 1 | 54.75 | 0.01 | 29.10 | 0.34 | 0.01 | 0.07 | 11.12 | 3.99 | 0.21 | 0.00 | 0.01 | 99.61 | 2.15 | 40.88 | 56.97 |
| Gh | 2 | 55.15 | 0.03 | 29.16 | 0.35 | 0.02 | 0.01 | 11.09 | 4.31 | 0.22 | 0.00 | 0.00 | 100.34 | 2.18 | 42.78 | 55.04 |
| Gh | 3 | 55.24 | 0.02 | 28.94 | 0.31 | 0.00 | 0.00 | 10.65 | 4.21 | 0.21 | 0.02 | 0.01 | 99.61 | 2.15 | 43.20 | 54.64 |
| Gh | 4 | 55.55 | 0.00 | 28.25 | 0.26 | 0.00 | 0.00 | 12.06 | 4.46 | 0.19 | 0.01 | 0.01 | 100.79 | 2.40 | 46.35 | 51.25 |
| Gh | 5 | 56.41 | 0.04 | 28.40 | 0.31 | 0.00 | 0.02 | 9.88 | 4.71 | 0.27 | 0.02 | 0.01 | 100.07 | 2.72 | 47.48 | 49.80 |
| Gh | 6 | 46.18 | 0.00 | 34.34 | 0.35 | 0.00 | 0.13 | 17.47 | 1.80 | 0.05 | 0.00 | 0.02 | 100.34 | 0.47 | 17.01 | 82.52 |
| Gh | 7 | 50.48 | 0.01 | 31.50 | 0.33 | 0.00 | 0.02 | 13.11 | 4.58 | 0.15 | 0.00 | 0.01 | 100.19 | 1.33 | 40.58 | 58.09 |
| Gh | 8 | 44.58 | 0.02 | 35.80 | 0.48 | 0.02 | 0.00 | 17.13 | 1.66 | 0.05 | 0.00 | 0.00 | 99.74 | 0.49 | 16.16 | 83.36 |
| Gh | 9 | 57.72 | 0.04 | 26.96 | 0.22 | 0.00 | 0.00 | 8.71 | 4.65 | 0.31 | 0.00 | 0.03 | 98.64 | 7.13 | 16.15 | 76.72 |
| Gh | 10 | 56.72 | 0.05 | 26.72 | 0.47 | 0.01 | 0.06 | 10.13 | 4.20 | 0.11 | 0.02 | 0.02 | 98.51 | 0.47 | 17.01 | 82.52 |
| Gh | 11 | 63.06 | 0.00 | 20.97 | 0.18 | 0.00 | 0.00 | 8.07 | 5.90 | 0.21 | 0.00 | 0.01 | 98.40 | 1.33 | 40.58 | 58.09 |
| G1 | | 49.78 | 1.13 | 5.81 | 13.34 | 0.55 | 14.60 | 11.57 | 1.29 | 0.57 | 0.00 | 0.02 | 98.66 | 7.46 | 16.87 | 75.67 |
| G1 | я | 49.50 | 1.34 | 6.03 | 13.40 | 0.61 | 14.06 | 11.36 | 1.37 | 0.63 | 0.00 | 0.02 | 98.32 | 8.20 | 17.84 | 73.96 |
| G1 | ц. | 62.04 | 0.00 | 23.59 | 0.13 | 0.01 | 0.00 | 5.75 | 7.12 | 0.35 | 0.01 | 0.00 | 99.00 | 1.14 | 30.70 | 68.16 |
| G1 | e to | 56.09 | 0.01 | 27.74 | 0.33 | 0.00 | 0.04 | 10.88 | 3.91 | 0.22 | 0.01 | 0.01 | 99.24 | 3.04 | 42.48 | 54.48 |
| G1 | Ore | 38.34 | 3.95 | 13.75 | 17.28 | 0.26 | 12.96 | 0.01 | 0.00 | 9.74 | 0.00 | 0.00 | 96.29 | 2.44 | 49.39 | 48.17 |
| G1 | 0 | 51.80 | 0.00 | 32.35 | 0.27 | 0.01 | 0.01 | 12.22 | 2.51 | 0.10 | 0.02 | 0.00 | 99.29 | 3.10 | 50.91 | 45.99 |
| G1 | | 54.24 | 0.00 | 28.39 | 0.25 | 0.00 | 0.00 | 13.17 | 3.67 | 0.19 | 0.02 | 0.01 | 99.94 | 4.32 | 58.16 | 37.52 |
| G2 | | 48.77 | 1.25 | 6.85 | 13.14 | 0.50 | 14.82 | 11.29 | 0.68 | 0.64 | 0.00 | 0.00 | 97.94 | 9.19 | 9.76 | 81.05 |
| G2 | | 49.88 | 1.17 | 5.63 | 12.73 | 0.51 | 14.75 | 11.40 | 1.20 | 0.53 | 0.00 | 0.00 | 97.80 | 7.13 | 16.15 | 76.72 |
| G2 | _ | 52.15 | 0.02 | 31.20 | 0.35 | 0.00 | 0.04 | 13.65 | 3.11 | 0.17 | 0.00 | 0.02 | 100.71 | 1.68 | 30.78 | 67.54 |
| G2 | nin | 56.27 | 0.01 | 28.86 | 0.34 | 0.01 | 0.05 | 10.48 | 4.44 | 0.22 | 0.01 | 0.04 | 100.73 | 2.22 | 44.85 | 52.93 |
| G2 | to | 56.51 | 0.01 | 28.54 | 0.35 | 0.01 | 0.05 | 10.19 | 4.21 | 0.24 | 0.01 | 0.01 | 100.13 | 2.51 | 44.11 | 53.38 |
| G2 | ore | 52.76 | 0.02 | 30.69 | 0.35 | 0.01 | 0.00 | 13.10 | 2.95 | 0.11 | 0.00 | 0.04 | 100.03 | 1.14 | 30.70 | 68.16 |
| G2 | Ŭ | 57.21 | 0.02 | 28.38 | 0.31 | 0.01 | 0.00 | 9.70 | 4.60 | 0.26 | 0.00 | 0.05 | 100.54 | 2.68 | 47.37 | 49.95 |
| G2 | | 57.76 | 0.02 | 27.34 | 0.33 | 0.00 | 0.07 | 8 89 | 4 92 | 0.30 | 0.00 | 0.00 | 99.63 | 3 10 | 50.91 | 45 99 |
| G2 | | 58.12 | 0.01 | 27.42 | 0.25 | 0.00 | 0.00 | 8 80 | 5.21 | 0.31 | 0.00 | 0.00 | 100.12 | 3 13 | 52 52 | 44 35 |
| G3 | | 55 79 | 0.01 | 27.12 | 0.26 | 0.00 | 0.00 | 10.52 | 4 36 | 0.25 | 0.00 | 0.00 | 99.04 | 1 99 | 39.21 | 58.81 |
| G3 | я | 56.28 | 0.00 | 20.33 | 0.10 | 0.00 | 0.00 | 10.12 | 1.50 | 0.23 | 0.02 | 0.00 | 100 71 | 2 34 | 46.13 | 51.53 |
| G3 | Ē | 56.23 | 0.00 | 29.55 | 0.17 | 0.01 | 0.00 | 10.12 | 4.55 | 0.23 | 0.02 | 0.00 | 100.71 | 2.54 | 46.15 | 51.55 |
| G3 | e to | 56.88 | 0.01 | 29.14 | 0.24 | 0.01 | 0.00 | 9.46 | 4.05 | 0.24 | 0.02 | 0.02 | 100.70 | 2.40 | 10.30 | 48.17 |
| 63 | Cor | 56.51 | 0.01 | 20.71 | 0.20 | 0.00 | 0.08 | 0.04 | 4.05 | 0.24 | 0.00 | 0.03 | 100.34 | 2.44 | 50.26 | 47.02 |
| G3 | 0 | 60.29 | 0.02 | 29.23 | 0.20 | 0.00 | 0.08 | 7.04 | 5.65 | 0.23 | 0.03 | 0.02 | 100.50 | 4.20 | 50.50 | 47.05 |
| G4 | | 55 27 | 0.02 | 20.03 | 0.21 | 0.00 | 0.03 | 10.29 | 5.05 4.02 | 0.42 | 0.00 | 0.00 | 07.07 | 4.52 | 20.79 | 57.52 |
| 04 C4 | | 53.27 | 0.00 | 20.94 | 0.16 | 0.00 | 0.04 | 10.20 | 4.05 | 0.27 | 0.05 | 0.05 | 97.07 | 1.00 | 20.21 | 50.01 |
| G4 | | 54.08 | 0.02 | 29.04 | 0.30 | 0.01 | 0.03 | 11.85 | 3.95 | 0.20 | 0.01 | 0.00 | 100.15 | 1.99 | 39.21 | 54.40 |
| G4 | В | 56.05 | 0.05 | 28.48 | 0.58 | 0.00 | 0.04 | 10.05 | 3.91 | 0.28 | 0.03 | 0.01 | 99.22 | 3.04 | 42.48 | 52.02 |
| G4 | 0 10 | 56.24 | 0.01 | 27.50 | 0.17 | 0.00 | 0.02 | 10.25 | 4.09 | 0.24 | 0.01 | 0.00 | 98.59 | 2.22 | 44.85 | 52.93 |
| G4 | re t | 56.57 | 0.01 | 27.65 | 0.29 | 0.01 | 0.00 | 11.59 | 4.21 | 0.23 | 0.00 | 0.00 | 100.56 | 2.60 | 50.36 | 4/.03 |
| G4 | Co | 54.77 | 0.01 | 28.06 | 0.28 | 0.01 | 0.00 | 11.46 | 3.78 | 0.20 | 0.02 | 0.03 | 98.62 | 2.15 | 40.88 | 56.97 |
| G4 | - | 63.36 | 0.05 | 24.34 | 0.21 | 0.00 | 0.05 | 5.11 | 7.35 | 0.31 | 0.00 | 0.04 | 100.82 | 2.15 | 43.20 | 54.64 |
| G4 | | 53.56 | 0.02 | 28.87 | 0.38 | 0.01 | 0.00 | 11.70 | 3.63 | 0.21 | 0.00 | 0.02 | 98.40 | 2.72 | 47.48 | 49.80 |
| G4 | | 65.68 | 0.01 | 22.50 | 0.19 | 0.00 | 0.01 | 3.82 | 6.93 | 0.65 | 0.00 | 0.00 | 99.79 | 4.23 | 65.53 | 30.25 |

| Table 2 Representat | ive electron micro | probe analyses o | of plagioclase from | m the Ghaha | n granitoids († | wt%) |
|-------------------------|--------------------|------------------|---------------------|-------------|-----------------|------|
| - able - representation | | | | | | |

nucleation density at a point where L reaches 0. Nucleation rate J is calculated through $n_0=J/G$ in which n_0 is a constant quantity which is obtained from the intersection of the cumulative density axis of crystal size.

This equation is equal to:

 $Tr = (-1/G \times m)/31,536,000$

where Tr is the residence time (years), G is the growth rate of crystals (mm/s), m is the slope of the CSD curve of the log-linear data, and the 31,536,000 coefficient converts seconds to years.

4 Results

4.1 Whole-rock chemistry

The bulk-rock chemical data (Table 4) were plotted in a variety of standard petrological diagrams (Fig. 5). Accordingly, the Ghahan/Sarbadan suite is characterized as diorite or granodiorite (Fig. 5a), calc-alkalic (Fig. 5b), metaluminous (Fig. 5c), I-type (Fig. 5d), with a volcanic arc affinity (Fig. 5e). In a diagram of mantle-normalized (Fig. 5f) abundances of selected major and trace elements,

data from the two localities are slightly offset, but track quite closely.

4.2 Mineral chemistry

4.2.1 Amphibole

Amphiboles from the Ghahan and Sarbadan localities show a narrow range of compositional variation. Structural formulas, calculated on an anhydrous basis assuming 23 oxygen atoms per half unit cell, place them in the calcic group (magnesio-hornblende) according to the classification scheme of Leake et al. (2004) (Fig. 6a). In a (Na + Ca + K) vs. Si diagram, which discriminates between metamorphic and igneous amphiboles, all data fall decisively in the igneous field, (Fig. 6b). In the TiO₂ vs. Al₂O₃ diagram, data for all the amphibole crystals plot in the crust-mantle mixed source field (Fig. 6c).

4.2.2 Plagioclase

Plagioclase, the most common phase, mainly occurs as an isolated euhedral to subhedral tabular crystal. The anorthite component decreases from diorite (An_{30-83}) to

 Table 3 Representative electron microprobe analyses of plagioclase from the Sarbadan granitoids (wt%)

| Sample | No | SiO | TiO | AlaOa | FeO | MnO | MσO | CaO | NacO | K ₂ O | NiO | PaOs | Total | Or | Δb | Δn |
|----------|-------|-------|------|-------|------|------|-------|--------------|------|------------------|------|------|--------|--------------|----------------|----------------|
| Sa | 1 | 54.45 | 0.02 | 28.51 | 0.26 | 0.02 | 0.19 | 9.54 | 7 57 | 0.26 | 0.00 | 0.04 | 100.86 | 2.06 | 60.08 | 37.86 |
| Sa | 2 | 60.39 | 0.02 | 20.31 | 0.20 | 0.02 | 0.15 | 6.91 | 8 32 | 0.20 | 0.00 | 0.04 | 100.80 | 2.00 4 54 | 67.45 | 28.01 |
| Sa | 3 | 52 55 | 0.04 | 30.16 | 0.15 | 0.00 | 0.05 | 10.64 | 6.98 | 0.21 | 0.02 | 0.00 | 100.81 | 1.68 | 55.80 | 42 53 |
| Sa | 4 | 53.36 | 0.00 | 29.70 | 0.27 | 0.00 | 0.00 | 10.98 | 5.71 | 0.25 | 0.01 | 0.02 | 100.30 | 2.18 | 49.87 | 47.95 |
| Sa | 5 | 55.88 | 0.02 | 27.34 | 0.26 | 0.00 | 0.08 | 8.13 | 8.76 | 0.33 | 0.01 | 0.00 | 100.81 | 2.51 | 66.59 | 30.90 |
| Sa | 6 | 62.73 | 0.02 | 22.63 | 0.19 | 0.00 | 0.00 | 5.76 | 6.40 | 0.35 | 0.00 | 0.02 | 98.10 | 2.28 | 53.53 | 44.19 |
| Sa | 7 | 59.83 | 0.04 | 24.88 | 0.31 | 0.01 | 0.02 | 8.85 | 4.16 | 0.25 | 0.00 | 0.00 | 98.35 | 3.29 | 62.11 | 34.60 |
| Sa | 8 | 56.77 | 0.00 | 26.61 | 0.34 | 0.00 | 0.01 | 10.43 | 4.58 | 0.30 | 0.00 | 0.01 | 99.05 | 2.80 | 59.87 | 37.32 |
| Sa | 9 | 65.97 | 0.05 | 18.85 | 0.09 | 0.00 | 0.00 | 0.07 | 1.07 | 13.01 | 0.01 | 0.01 | 99.13 | 3.15 | 62.88 | 33.97 |
| Sa | 10 | 57.15 | 0.05 | 27.48 | 0.14 | 0.01 | 0.00 | 9.79 | 4.73 | 0.19 | 0.00 | 0.02 | 99.56 | 2.44 | 52.26 | 45.30 |
| Sa | 11 | 66.73 | 0.03 | 18.63 | 0.14 | 0.02 | 0.00 | 0.35 | 1.80 | 10.80 | 0.00 | 0.00 | 98.50 | 2.06 | 60.09 | 37.85 |
| Sa | 12 | 66.38 | 0.00 | 19.53 | 0.06 | 0.00 | 0.00 | 0.01 | 0.31 | 12.68 | 0.00 | 0.00 | 98.97 | 1.75 | 56.78 | 41.47 |
| Sa | 13 | 55.46 | 0.07 | 25.05 | 0.39 | 0.01 | 0.02 | 9.12 | 4.31 | 0.22 | 0.00 | 0.06 | 94.71 | 2.36 | 65.52 | 32.11 |
| Sa | 14 | 60.96 | 0.07 | 23.51 | 0.21 | 0.01 | 0.00 | 9.55 | 4.41 | 0.26 | 0.01 | 0.00 | 98.99 | 2.51 | 50.89 | 46.61 |
| Sa | 15 | 65.98 | 0.04 | 18./1 | 0.10 | 0.00 | 0.01 | 0.03 | 0.73 | 13.98 | 0.00 | 0.00 | 99.58 | 3.02 | 65.26 | 31.72 |
| Sa | 10 | 58.03 | 0.00 | 27.32 | 0.19 | 0.01 | 0.01 | 10.13 | 6.12 | 0.31 | 0.01 | 0.04 | 100.17 | 1./1 | 55.46 | 46.95 |
| Sa Sa | 17 | 58.04 | 0.01 | 20.20 | 0.17 | 0.00 | 0.00 | 8.70 | 7.05 | 0.42 | 0.00 | 0.00 | 100.00 | 2.42 | 58.86 | 42.12 38.66 |
| Sa Sa | 10 | 57 54 | 0.01 | 26.19 | 0.10 | 0.00 | 0.00 | 8.38 | 6 79 | 0.45 | 0.00 | 0.02 | 99.29 | 2.47 | 58.80 | 39.12 |
| Sa | 20 | 58.96 | 0.00 | 25.84 | 0.20 | 0.00 | 0.00 | 7 99 | 7 45 | 0.30 | 0.00 | 0.00 | 100.87 | 2.05 | 61 39 | 36.38 |
| Sa | 20 | 57.10 | 0.02 | 26.80 | 0.18 | 0.01 | 0.00 | 9.60 | 5.93 | 0.33 | 0.01 | 0.00 | 100.07 | 1.90 | 51 78 | 46.32 |
| Sa | 22 | 56.70 | 0.02 | 27.34 | 0.15 | 0.00 | 0.01 | 10.01 | 6.47 | 0.25 | 0.00 | 0.04 | 100.99 | 1.35 | 53.18 | 45.47 |
| Sa | 23 | 55.47 | 0.00 | 28.35 | 0.18 | 0.00 | 0.00 | 10.70 | 5.90 | 0.26 | 0.00 | 0.00 | 100.86 | 1.43 | 49.23 | 49.34 |
| Sa | 24 | 56.55 | 0.01 | 26.19 | 0.18 | 0.01 | 0.00 | 9.25 | 7.77 | 0.38 | 0.00 | 0.02 | 100.36 | 1.90 | 59.17 | 38.93 |
| Sa | 25 | 55.99 | 0.00 | 26.97 | 0.19 | 0.00 | 0.00 | 9.85 | 7.49 | 0.33 | 0.02 | 0.01 | 100.85 | 1.65 | 56.96 | 41.39 |
| Sa | 26 | 59.03 | 0.01 | 25.62 | 0.21 | 0.00 | 0.02 | 8.18 | 7.32 | 0.43 | 0.00 | 0.03 | 100.85 | 2.33 | 60.38 | 37.29 |
| Sa | 27 | 58.73 | 0.01 | 27.46 | 0.20 | 0.00 | 0.00 | 9.61 | 4.47 | 0.27 | 0.02 | 0.00 | 100.77 | 3.84 | 62.12 | 34.04 |
| Sa | 28 | 59.13 | 0.01 | 27.17 | 0.19 | 0.02 | 0.00 | 9.07 | 5.04 | 0.29 | 0.00 | 0.01 | 100.93 | 3.28 | 65.57 | 31.15 |
| Sa | 29 | 61.66 | 0.01 | 23.90 | 0.14 | 0.00 | 0.00 | 5.89 | 8.11 | 0.52 | 0.00 | 0.00 | 100.23 | 2.92 | 69.28 | 27.80 |
| Sa | 30 | 57.55 | 0.01 | 26.72 | 0.17 | 0.01 | 0.00 | 9.04 | 6.91 | 0.35 | 0.01 | 0.03 | 100.80 | 1.90 | 56.94 | 41.16 |
| G5 | | 57.81 | 0.03 | 27.53 | 0.30 | 0.01 | 0.12 | 9.91 | 4.63 | 0.28 | 0.00 | 0.02 | 100.64 | 2.84 | 46.93 | 50.23 |
| G5 | | 57.50 | 0.02 | 28.14 | 0.31 | 0.01 | 0.10 | 9.48 | 4.71 | 0.26 | 0.00 | 0.05 | 100.58 | 2.68 | 48.51 | 48.82 |
| GS | ш | 56.96 | 0.00 | 28.14 | 0.31 | 0.00 | 0.14 | 9.49 | 5.10 | 0.30 | 0.00 | 0.00 | 100.44 | 2.96 | 50.27 | 46.77 |
| 65 | 0 11 | 54.81 | 0.00 | 27.67 | 0.19 | 0.01 | 0.01 | 10.28 | 5.93 | 0.30 | 0.00 | 0.04 | 99.24 | 2.64 | 52.15 | 45.21 |
| 65 | re t | 56.84 | 0.00 | 25.17 | 0.15 | 0.01 | 0.01 | 9.59 | 0.78 | 0.31 | 0.01 | 0.00 | 98.87 | 2.01 | 57.05 | 40.34 |
| 65 | Co | 56.59 | 0.04 | 25.99 | 0.19 | 0.00 | 0.00 | 8.93 | 6.81 | 0.45 | 0.00 | 0.00 | 99.00 | 3.84 | 58.08 | 38.08 |
| 65 | | 57.64 | 0.02 | 26.18 | 0.19 | 0.00 | 0.01 | 8.84 | 0.00 | 0.41 | 0.00 | 0.02 | 99.10 | 3.5/ | 57.90 | 38.4/ |
| G5 | | 62.24 | 0.00 | 25.87 | 0.10 | 0.00 | 0.01 | 8.30 5.94 | 6.30 | 0.39 | 0.01 | 0.00 | 99.20 | 3.47 | 58.41 | 28.11 |
| <u> </u> | | 56.02 | 0.03 | 23.02 | 0.20 | 0.00 | 0.00 | 0.04 | 0.92 | 0.42 | 0.00 | 0.01 | 00.67 | 4.09 | 47.05 | 20.40 |
| G0 C6 | ч | 57.17 | 0.01 | 28.40 | 0.32 | 0.00 | 0.06 | 9.94 | 4.05 | 0.24 | 0.00 | 0.05 | 99.07 | 2.44 | 47.05 | 30.31 46.00 |
| 00 G6 | nir | 58.56 | 0.02 | 20.34 | 0.30 | 0.00 | 0.07 | 9.30 | 5.01 | 0.20 | 0.02 | 0.01 | 100.00 | 2.01 | 54.02 | 40.99 |
| 00 G6 | e to | 58.30 | 0.00 | 27.33 | 0.34 | 0.00 | 0.03 | 8.57 | 5.01 | 0.32 | 0.00 | 0.03 | 100.83 | 2 55 | 55.00 | 41.95 |
| 00 G6 | Con | 62 50 | 0.00 | 21.47 | 0.20 | 0.00 | 0.12 | 5.40 | 6.81 | 0.57 | 0.00 | 0.00 | 100.92 | 5.55 | 66.86 | 27.64 |
| G6 |) | 63.42 | 0.00 | 24.00 | 0.18 | 0.00 | 0.04 | 4 86 | 7.58 | 0.30 | 0.00 | 0.02 | 100.00 | 1 48 | 72 33 | 27.04 |
| G7 | | 56.81 | 0.01 | 28.14 | 0.32 | 0.00 | 0.18 | 9.50 | 4 92 | 0.76 | 0.02 | 0.00 | 100.14 | 2.62 | 49 55 | 47.83 |
| G7 | rim | 58.20 | 0.01 | 27 29 | 0.31 | 0.00 | 0.11 | 8 90 | 5 43 | 0.20 | 0.02 | 0.00 | 100.14 | 3.04 | 53 29 | 43.67 |
| G7 | to i | 57.61 | 0.00 | 28.18 | 0.30 | 0.01 | 0.04 | 9.26 | 5.13 | 0.32 | 0.01 | 0.00 | 100.86 | 3.84 | 58.08 | 38.08 |
| G7 | ore | 57.43 | 0.00 | 28.16 | 0.29 | 0.00 | 0.05 | 9.30 | 5.19 | 0.30 | 0.01 | 0.00 | 100.73 | 2.64 | 52.15 | 45.21 |
| G7 | Ŭ | 56.21 | 0.03 | 28.76 | 0.35 | 0.00 | 0.12 | 9.93 | 4 95 | 0.26 | 0.00 | 0.01 | 100.62 | 3 47 | 58 41 | 38.11 |
| G8 | | 57.82 | 0.02 | 27.63 | 0.21 | 0.00 | 0.05 | 9.63 | 4.95 | 0.26 | 0.00 | 0.02 | 100.59 | 2.68 | 47.37 | 49.95 |
| G8 | | 58.27 | 0.01 | 27.12 | 0.25 | 0.00 | 0.01 | 10.02 | 5.09 | 0.29 | 0.00 | 0.00 | 101.06 | 3 13 | 52 52 | 44 35 |
| C8 | | 59.27 | 0.00 | 26.82 | 0.20 | 0.00 | 0.14 | 7 71 | 5.07 | 0.20 | 0.00 | 0.00 | 100.82 | 8 20 | 27.84 | 63.06 |
| 68 | | 60.06 | 0.00 | 20.02 | 0.20 | 0.00 | 0.14 | 8.02 | 5 17 | 0.30 | 0.01 | 0.00 | 100.02 | 2 84 | ∠7.04 16.02 | 50.22 |
| 00 | | 57.10 | 0.01 | 20.23 | 0.20 | 0.00 | 0.01 | 0.02 | 5.17 | 0.34 | 0.00 | 0.00 | 100.94 | 2.04 | 40.93 | 40.00 |
| 68 | Е | 57.19 | 0.04 | 21.42 | 0.22 | 0.02 | 0.00 | 9.76 | 5.09 | 0.30 | 0.00 | 0.02 | 100.06 | 2.68 | 48.51 | 48.82 |
| G8 | in c | 58.52 | 0.00 | 26.53 | 0.27 | 0.00 | 0.02 | 9.47 | 4.85 | 0.37 | 0.00 | 0.00 | 100.03 | 2.96 | 50.27 | 46.77 |
| G8 | re tu | 59.42 | 0.01 | 25.35 | 0.17 | 0.00 | 0.00 | 7.54 | 6.81 | 0.52 | 0.00 | 0.05 | 99.87 | 3.02 | 60.17 | 36.81 |
| G8 | Col | 57.21 | 0.02 | 26.84 | 0.19 | 0.01 | 0.00 | 9.33 | 6.78 | 0.36 | 0.02 | 0.01 | 100.77 | 1.95 | 55.70 | 42.36 |
| G8 | - | 57.65 | 0.01 | 26.05 | 0.17 | 0.01 | 0.00 | 8.65 | 6.86 | 0.46 | 0.00 | 0.00 | 99.86 | 2.53 | 57.44 | 40.02 |
| G8 | | 56.28 | 0.03 | 26.23 | 0.20 | 0.01 | 0.00 | 9.12 | 7.67 | 0.35 | 0.00 | 0.02 | 99.91 | 1.78 | 59.27 | 38.95 |
| G8 | | 58.09 | 0.03 | 26.05 | 0.17 | 0.00 | 0.00 | 8.34 | 7.49 | 0.33 | 0.00 | 0.01 | 100.51 | 1.76 | 60.82 | 37.42 |
| G8 | | 59.42 | 0.00 | 25.59 | 0.17 | 0.00 | -0.01 | 7.70 | 7.78 | 0.40 | 0.02 | 0.01 | 101.08 | 2.14 | 63.26 | 34.60 |
| G8 | | 59.69 | 0.01 | 26.23 | 0.26 | 0.00 | 0.01 | 8.78 | 4.71 | 0.39 | 0.02 | 0.02 | 100.12 | 4.48 | 63.17 | 32.35 |

granodiorite (An_{23-64}) . Geochemically, they can be classified as predominantly andesine and to a lesser degree

oligoclase and labradorite. EPMA profiles of An content vs. distance (Fig. 7 indicates both oscillatory and normal

| Rock type | | Granodiorite | | | Diorite | |
|--------------------------------|----------|--------------|----------|----------|---------|----------|
| Sample | S1 | S2 | S3 | G1 | G2 | G3 |
| SiO ₂ | 66.06 | 66.24 | 65.60 | 62.55 | 55.02 | 61.79 |
| TiO ₂ | 0.45 | 0.43 | 0.57 | 0.57 | 0.71 | 0.58 |
| Al_2O_3 | 15.87 | 15.55 | 15.81 | 16.73 | 18.75 | 16.33 |
| Fe ₂ O ₃ | 3.79 | 4.17 | 5.13 | 5.63 | 4.60 | 5.81 |
| MnO | 0.04 | 0.06 | 0.14 | 0.10 | 0.10 | 0.09 |
| MgO | 1.55 | 1.71 | 2.94 | 2.21 | 4.01 | 2.31 |
| CaO | 4.86 | 4.18 | 5.22 | 5.67 | 8.91 | 5.65 |
| Na ₂ O | 4.31 | 3.77 | 2.86 | 3.54 | 3.38 | 3.62 |
| K ₂ O | 2.00 | 2.44 | 1.57 | 1.82 | 0.57 | 2.23 |
| P_2O_5 | 0.09 | 0.09 | 0.08 | 0.12 | 0.12 | 0.12 |
| LoI | 0.32 | 0.65 | 0.57 | 1.43 | 1.12 | 1.35 |
| Total | 99.34 | 99.27 | 100.49 | 100.37 | 97.29 | 99.89 |
| Cs | 0.55 | 0.48 | 0.60 | 0.50 | 1.59 | 0.18 |
| Rb | 33.00 | 39.45 | 42.68 | 63.21 | 60.11 | 19.11 |
| Ba | 550.04 | 691.66 | 666.43 | 600.00 | 470.59 | 409.89 |
| Th | 10.90 | 9.05 | 9.06 | 4.31 | 5.25 | 3.94 |
| U | 2.74 | 2.77 | 2.74 | 1.53 | 1.51 | 1.29 |
| Nb | 5.38 | 4.97 | 5.02 | 4.50 | 5.74 | 2.86 |
| Та | 0.60 | 0.57 | 0.62 | 0.64 | 0.55 | 0.18 |
| K | 16596.72 | 20232.68 | 13064.20 | 15108.92 | 4695.21 | 18536.78 |
| La | 11.96 | 14.71 | 12.77 | 13.00 | 24.24 | 6.05 |
| Ce | 20.09 | 25.37 | 22.68 | 32.00 | 44.41 | 12.86 |
| Pb | 4.06 | 6.87 | 6.62 | 5.00 | 7.34 | 3.75 |
| Sr | 356.03 | 327.24 | 304.40 | 308.50 | 253.15 | 410.77 |
| Р | 373.36 | 384.87 | 345.02 | 530.89 | 503.62 | 535.85 |
| Nd | 8.48 | 10.44 | 9.91 | 12.00 | 17.25 | 8.25 |
| Hf | 3.29 | 3.22 | 3.58 | 1.95 | 3.14 | 1.72 |
| Zr | 132.81 | 132.28 | 151.30 | 59.00 | 120.59 | 64.51 |
| Sm | 2.00 | 2.24 | 2.25 | 2.34 | 3.33 | 2.42 |
| Ti | 2687.55 | 2592.33 | 3422.09 | 3437.13 | 4245.96 | 3459.07 |
| Tb | 0.31 | 0.34 | 0.33 | 0.61 | 0.48 | 0.45 |
| Y | 12.09 | 13.13 | 13.10 | 12.90 | 17.40 | 17.04 |
| Er | 1.25 | 1.36 | 1.31 | 1.45 | 1.47 | 1.84 |
| Tm | 0.19 | 0.20 | 0.21 | 0.19 | 0.25 | 0.27 |
| Yb | 1.37 | 1.45 | 1.47 | 1.80 | 1.73 | 1.79 |
| Со | 34.62 | 43.23 | 44.50 | 13.80 | 33.90 | 10.13 |
| Cr | 9.62 | 10.81 | 9.91 | 9.00 | 24.39 | 21.69 |
| Cu | 6.94 | 6.15 | 19.35 | 7.00 | 6.73 | 1.87 |
| Dy | 1.89 | 2.02 | 2.03 | 3.10 | 2.58 | 2.90 |
| Mo | 0.33 | 0.44 | 0.54 | 0.40 | 0.80 | 0.51 |
| Ni | 9.45 | 9.36 | 10.05 | 2.00 | 64.10 | 22.00 |
| Sc | 8.40 | 8.56 | 8.32 | 63.60 | 11.41 | 20.22 |
| V | 82.35 | 83.11 | 74.74 | 110.00 | 108.61 | 175.30 |
| Zn | 10.92 | 69.44 | 12.21 | 83.00 | 40.84 | 58.87 |
| Ga | 13.61 | 13.65 | 13.43 | 19.01 | 16.24 | 16.39 |
| As | 20.25 | -0.20 | -0.17 | 9.70 | 0.49 | 4.75 |
| Sn | 0.88 | 0.68 | 0.58 | 1.00 | 0.90 | 0.52 |
| Sb | 1.73 | 0.16 | 0.12 | 0.45 | 0.90 | 1.23 |
| Pr | 2.24 | 2.81 | 2.59 | 2.99 | 4.79 | 1.81 |
| Eu | 0.71 | 0.83 | 0.78 | 0.83 | 0.91 | 0.76 |
| Gd | 1.97 | 2.30 | 2.19 | 1.95 | 3.22 | 2.65 |
| Но | 0.43 | 0.46 | 0.47 | 0.61 | 0.61 | 0.63 |
| Lu | 0.23 | 0.24 | 0.24 | 0.23 | 0.26 | 0.28 |

Table 4 Major (wt%), trace and rare earth element (ppm) data of representative rock samples from the Ghahan and Sarbadan granitoids

zoning). Some samples demonstrate high An variations, with the maximum difference in X_{An} in Tafresh plagioclase being more than 30 mol%.

4.3 Crystal Size Distribution (CSD)

4.3.1 3D shape

Table 5 presents the distribution of 2D length and width (intersection dimensions) data of 3,191 measured grains, and Table 6 contains calculated crystal shape data. The two-dimensional size average of plagioclase in the Tafresh granitoids is 0.45 millimeters. Crystal shapes are defined

in terms of 'aspect ratio', that is, the ratio of short: intermediate: long (S:I:L) dimensions. Fig. 8 shows frequency vs. 2D sa/la (short axis/long axis) histograms, together with the best-fit shape output by CSDSlice5 Spreadsheet (Morgan & Jerram 2006), which compares the distribution of 2D size measurements to a database of shape curves for random sections through different crystal shapes and determines a best fit 3D crystal habit based on regression calculations and fit to the database. Dotted lines show the most similar crystal shapes to our data, and have been considered to be similar to our sample crystal factors



Fig. 5. Chemical classification of rocks from the Ghahan and Sarbadan stocks. (a) SiO₂ vs. K₂O+Na₂O plot (Middlemost, 1994), composition of studied samples range from diorite (Ghahan stock) to granodiorite (Sarbadan stock); (b) AFM (A=Na₂O+K₂O, F=FeOt, M=MgO) diagram with differentiation lines of Irvine and Barager (1971), showing a calc-alkaline affinity for the granitoids; (c) classification of the Ghahan and Sarbadan stocks in A/CNK vs. A/NK diagram [ANK=molar Al₂O₃/(Na₂O+K₂O) and ACNK=molar Al₂O₃/(CaO+Na₂O+K₂O)], (Shand, 1943); (d) K₂O vs. Na₂O classification diagram for discrimination of I- and S-type granitoids (Chappell and White, 2001), all samples fall in the field of I-type granitoids; (e) tectonic classification diagrams (Ta+Yb) vs. Rb (Pearce et al. 1984). WPG: within-plate granites; VAG: volcanic are granites; syn-COLG: syn-collision granites; ORG: oceanic ridge granites; (f) primitive mantle normalized multi-element diagram. Normalizing values are after (McDonough and Sun, 1995). The dark triangle (Ghahan samples) and dark squares (Sarbadan samples) are from Mirnejad et al. (2018).

(Fig. 8). The I/L vs. S/I diagram of Zingg (1935) is used here to present the data and to evaluate the threedimensional (3D) shapes (Fig. 9). In this figure, the plagioclase crystals occur as prolate to bladed in 3D shape. **4.3.2 Residence time of solidification**

The CSD curves for plagioclase from Tafresh granitoids are presented in Fig. 10. Most CSD curves do not follow a

classic straight line trend. Slopes of CSD curves for Tafresh plagioclase (Fig. 10) become the basis for calculating an average crystal residence time by employing the equation:

$Tr = (-1/G \times m)/31536000$

These regression lines are drawn based on all the size intervals, and since the relatively small crystals are more



Fig. 6. (a) Diagrams showing the classification of hornblendes according to the nomenclature of Leake et al. (1997); (b) in the (Na+Ca+K) vs. Si, all data fall decisively in the igneous field; (c) Plot of TiO₂ vs. Al₂O₃ for hornblende, all hornblende crystals fall in the crust-mantle mixed source field (Jiang and An, 1984).

frequent, these lines mostly follow the segments of CSD curves which are related to these crystals. Cashman (1993)

Table 5 The estimated dimension of short (x), medium (y) and length (z) axis in CSDSlice environment of plagioclase for the the Ghahan and Sarbadan granitoids

| | | | - | |
|--------|--------|-------|--------------|------|
| Sampla | Counto | | 3D Shape | |
| Sample | Counts | Short | Intermediate | Long |
| G1 | 332 | 1 | 1.9 | 2.8 |
| G2 | 424 | 1 | 1.3 | 2.3 |
| G3 | 531 | 1 | 1.4 | 2.5 |
| G4 | 339 | 1 | 1.8 | 3.2 |
| G5 | 354 | 1 | 1.9 | 2.5 |
| G6 | 473 | 1 | 1.5 | 2.2 |
| G7 | 373 | 1 | 1.4 | 2.6 |
| G8 | 365 | 1 | 1.4 | 2.5 |
| Total | 3191 | 1 | 1.7 | 2.5 |
| | | | | |

Table 6 Line slope, primary nucleation density, and calculated of plagioclase crystallization from the Ghahan and Sarbadan granitoids

| Samula | m= | $n_0 =$ | Tr= |
|--------|------------------|----------------------|-------------------|
| Sample | regression slope | regression intercept | (-1/G×m)/31536000 |
| G1 | -1.07 | -0.41 | 339.29 |
| G2 | -1.31 | 0.35 | 415.40 |
| G3 | -1.9 | 1.92 | 602.49 |
| G4 | -1.32 | 0.6 | 418.57 |
| G5 | -1.71 | 1.16 | 542.24 |
| G6 | -2.06 | 1.74 | 653.22 |
| G7 | -1.2 | -0.09 | 380.52 |
| G8 | -1.12 | -0.26 | 355.15 |
| Total | -1.46 | 1.82 | 462.96 |

noted an important relationship between the cooling rate, and correspondingly, the crystallization rate, with the rate of the plagioclase crystal growth. For example, for a cooling interval of 3 years, which might pertain to a dike, the crystal growth rate is 10^{-9} mm/s, whereas for a cooling interval of 300 years, the growth rate is only one order of magnitude different (10^{-10} mm/s). Investigations by Garrido et al. (2001) and Cheng (2013) of igneous bodies emplaced in the shallow crust similarly indicate a growth rate of ~ 10^{-10} mm. Cashman (1993) suggested a residence period of less than 1,000 years for the final ascent of shallow intrusive bodies' chambers near the surface and related to the final stages of magma, such as those in the Tafresh granitoids. The crystal residence times of plagioclase from Tafresh are calculated to be less than 1,000 years (Table 6).

5 Discussion

5.1 Petrogenesis

As mentioned, Sarbadan and Ghahan granitoids exhibit Nb–Ta–Ti depletion (Fig. 5f), and relative enrichment in Ba, Th, Rb, U, Pb, and K. High Rb, Th, and K, and low Sr, P, and Ti values are diagnostic of typical crustal melts (Harris et al., 1986; Chappell White et al., 2001), suggesting some continental crustal involvement in development of the igneous province. Either interaction occurred between crustal materials or mafic magmas during ascent, or the magma came from an alreadyenriched mantle source. The common occurrence of magnetite and hornblende in Tafresh indicates a high oxidation state and elevated water content of the parent



Fig. 7. Measured EMPA profiles for the analyzed plagioclase phenocrysts of the Ghahan granitoids (Samples G1 to G8).

magmas (Hanson, 1980; Saunders et al., 1980), and likely explains the lack of a distinct Eu negative anomaly. In general, the Nb-Ta trough is considered to be a typical characteristic of arc-related I-type granitoids (e.g. Rogers and Hawkesworth, 1989; Sajona et al., 1996). Variations in the enrichment of LREE could be attributed either to varying source compositions or to the amount of partial melting (Langmuir et al., 1977; Le Roex et al., 1987; 1989; Natland, 1989; Niu et al., 1996, 1999; Kamenetsky et al., 2000). On the tectonic discrimination diagram of Pearce et al., (1984), all sample compositions fall in the field of volcanic arc granitoids (Fig. 5e). High Rb, Th, K and low Sr, P and Ti values are compatible with typical crustal melts (Harris et al., 1986; Chappell and White, 1992) and suggest some upper crustal contamination during magmatic evolution.

High LREE and slightly concave-upward pattern MREE to HREE (Fig. 5f) can also reflect amphibole fractionation.



Fig. 8. (a–d) The processed rock sample in the Ghahan granitoids (Sample G1 to G4); (e–h) the processed rock sample in Sarbadan granitoids (Sample G5 to G8). Mosaic thin sections, binary images and two dimensional width/length distributions for each sample (red line) are shown along with their best fit crystal habit (dashed lines) as determined by CSDSlice (Morgan and Jerram, 2006). The numbers on the left corner of the CSDSlice diagram represent the aspect ratios as given by short axis: intermediate axis: long axis of the crystal. The outlines are plagioclase.



Fig. 9. The position of the average aspect ratio of plagioclase in granitoid rocks on the I/L vs. S/I diagram (Zingg, 1935). S, I and L represent small, intermediate and long axis, respectively.

Amphibole typically has a high distribution constant (KD) for Nb. The geochemical characteristics of the studied samples strongly suggest that rather high pressure fractional crystallization (involving amphibole but not garnet) is responsible here for the generation of this rock type. If garnet, which sequesters HREEs, had been involved in magma fractionation, abundances of HREEs would have been far lower than the observed abundances (Fujimaki et al., 1984). This indicates that Tafresh magma was not separated from a source melt in which garnet had been restite. Substantial amounts of garnet were not involved, whether as residue during partial melting, or during fractionation or final re-equilibration of the magma. The magmas were produced in a T-P-XH₂O environment outside the garnet stability field, possibly in an amphibolebearing mantle source.



Fig. 10. (a–i) CSD diagrams of the studied plagioclase crystals from Tafresh granitoids. Note the non-linear trend of the diagrams (especially in figs. b, h and i). The dotted lines are the regression lines calculated based on the frequency of crystal sizes, the horizontal axis unit is the millimeter.

5.2 Pressure and temperature conditions of crystallization

Aluminum-in-hornblende barometry has been widely used in calc-alkaline granitoids to calculate pressures of magmatic crystallization and corresponding emplacement depths (e.g., Hammarstrom and Zen, 1986; Hollister et al., 1987; Anderson and Smith, 1995; Stein and Dietl, 2001; Ernst, 2002), implying the computed pressure can be attributed to the crystallization of hornblende, and not necessarily to the consolidation of the bulk granitic rock (Ague, 1997). The theory of Al-hornblende barometry requires a buffer assemblage (quartz + plagioclase + alkali feldspar + biotite + hornblende + titanite + magnetite or ilmenite), which limits the applicability of the method (Leake and Said, 1994; Ague, 1997). It is standard practice to analyze Al in hornblende crystal rims, thus to represent the final (near-solidus) conditions of the magma. We have chosen the calibration of Schmidt (1992) as being more reliable than the other methods for considering the influence of temperature and oxygen fugacity on the pressure calculation (Stein and Dietl, 2001; Toummite et al., 2012). The Al-in-hornblende barometry of Schmidt (1992) is applied to calculate pressure and temperature (P $(\pm 0.6 \text{ kbar}) = -3.01 + 4.76 \text{ Altot, } r^2 = 0.99)$. In general, the lower Altot contents in the amphiboles of Tafresh granodiorites (with mean 0.95) are indicative of a low pressure of crystallization. Table 7 summarizes the results of crystallization pressures calculated from hornblende composition. Also, we used the $Al^{1\nu}+Al^{\nu}$ vs. Fe/(Fe+Mg) diagram of Anderson and Smith (1995) for determination of the amphibole crystallization pressure. It suggests that the amphiboles crystallized under pressures of about 3 kbar (Fig. 11a). Furthermore, in the diagram of Altor vs P (kbar) (Anderson and Smith, 1995), all of the samples indicate a crystallization temperature of less than 750°C (Fig. 11b) that are consistent with the calculated

temperatures and pressures (Table 7).

5.3 Estimation of oxygen fugacity (fO₂)

Oxygen fugacity (or oxygen potential) is a powerful mechanism for understanding the crystallization of igneous and metamorphic rocks under variable pressure, temperature, and melt composition conditions (Toummite et al., 2012). The intrinsic oxygen fugacity of a magma is related to its source material, which in turn is related to the tectonic setting. More highly oxidized magmas are commonly associated with convergent plate boundaries (Ewart, 1979), whereas felsic magmas formed by fractionation from mantle-derived magmas in rift zones are more reduced (Loiselle, 1979). Ridolfi et al. (2010) proposed an empirical universal amphibole sensor to estimate fO_2 , which, when applied to data for the Ghahan and Sarbadan rocks, provided calculated values of log fO_2 in the range of -10.1 to -13.3 (with a mean -12.52).

5.4 Magmatic processes

As noted, a CSD diagram is a quantitative description of relationships between crystal size and crystal population density (Cashman and Marsh, 1988; Higgins, 2000; 2002; 1998; 1994; 1996; Marsh, 1998; 1988). These diagrams show changes in frequency and crystal sizes as a function



Fig. 11. (a) The position of the Tafresh granitoids on a $Al^{V}+Al^{Vi}$ vs. Fe/(Fe+Mg) diagram (Anderson and Smit, 1995); (b) determination of the amphibole crystallization pressure, Al (t) vs. P (kbar) diagram (Anderson and Smit, 1995) suggests that the amphiboles crystallized under pressures of about 3 kbar.

| Table 7 | Crystallization | pressure | calculated | of amphibole |
|----------|-----------------|-----------|---------------|--------------|
| from the | Ghahan and S | arbadan g | granitoids (i | in kbar). |

| 0 | | Hammarstrom | Johnson & | Schmidt | Anderson | Continental |
|------|------|-------------|---------------|---------|----------|-------------|
| Sa | mple | & Zen 86 | Rutherford 89 | 92 | & Smith | depth (km) |
| | Gh1 | 1.4 | 1.0 | 2.0 | 2.0 | 3.3 |
| | Gh2 | 1.5 | 1.1 | 2.1 | 2.1 | 3.4 |
| | Gh3 | 1.4 | 1.0 | 2.0 | 2.0 | 3.3 |
| | Gh4 | 0.9 | 0.6 | 1.5 | 1.5 | 2.8 |
| | Gh5 | 1.0 | 0.7 | 1.7 | 1.7 | 3.0 |
| | Gh6 | 1.3 | 0.9 | 1.9 | 1.9 | 3.2 |
| | Gh7 | 1.0 | 0.7 | 1.7 | 1.7 | 3.0 |
| | Gh8 | 0.5 | 0.3 | 1.2 | 1.2 | 2.6 |
| | Gh9 | -0.1 | -0.2 | 0.6 | 0.6 | 2.2 |
| | Gh10 | -0.1 | -0.3 | 0.6 | 0.6 | 2.1 |
| | Gh11 | 2.3 | 1.8 | 2.9 | 2.8 | 4.3 |
| | Gh12 | 1.9 | 1.4 | 2.5 | 2.4 | * |
| | Gh13 | -0.2 | -0.3 | 0.5 | 0.5 | * |
| g | Gh14 | 0.8 | 0.5 | 1.5 | 1.4 | 2.8 |
| lha | Gh15 | 0.0 | -0.2 | 0.7 | 0.7 | * |
| Gha | Gh16 | -0.1 | -0.2 | 0.6 | 0.6 | * |
| 0 | Gh17 | -0.2 | -0.4 | 0.5 | 0.5 | * |
| | Gh18 | 1.1 | 0.8 | 1.8 | 1.7 | 3.1 |
| | Gh19 | 0.0 | -0.1 | 0.7 | 0.7 | * |
| | Gh20 | 1.1 | 0.7 | 1.7 | 1.7 | * |
| | Gh21 | 1.1 | 0.8 | 1.8 | 1.7 | * |
| | Gh22 | 0.9 | 0.6 | 1.5 | 1.5 | 2.9 |
| | Gh23 | 1.1 | 0.7 | 1.7 | 1.7 | 3.0 |
| | Gh24 | 0.8 | 0.5 | 1.4 | 1.4 | 2.8 |
| | Gh25 | 1.0 | 0.6 | 1.6 | 1.6 | 2.9 |
| | Gh26 | 1.7 | 1.3 | 2.3 | 2.3 | 3.6 |
| | Gh27 | 0.5 | 0.3 | 1.2 | 1.2 | 2.6 |
| | Sal | 0.7 | 0.4 | 1.4 | 1.3 | 2.7 |
| | Sa2 | 0.7 | 0.4 | 1.4 | 1.4 | 2.7 |
| | Sa3 | 1.5 | 1.1 | 2.1 | 2.1 | 3.4 |
| an | Sa4 | 1.2 | 0.9 | 1.9 | 1.9 | 3.2 |
| bad | Sa5 | 2.8 | 2.2 | 3.3 | 3.3 | * |
| Sarl | Sa6 | 0.7 | 0.4 | 1.4 | 1.4 | 2.7 |
| 01 | Sa7 | 0.7 | 0.4 | 1.4 | 1.4 | 2.7 |
| | Sa8 | 0.9 | 0.6 | 1.5 | 1.5 | 2.8 |
| | Sa9 | 0.4 | 0.2 | 1.1 | 1.1 | 2.5 |

of the latent residency time in the system. A linear CSD diagram indicates simple, monotonic cooling of magma, but if the diagram exhibits kinks, or changes of slope (Fig. 10B, C, H and I), additional processes must have been superimposed upon the cooling. One possibility is coarsening, whereby small crystals are consumed by annealing as large crystals grow larger (Fig. 12d). Another possibility is the mixing of populations of crystals that had formed in different sources during rapid ascent or magma mixing (Fig. 12e).

It is noteworthy that the Tafresh granitoids contain abundant globular or elliptical MMEs of various sizes, since the existence of MMEs in the Tafresh granitoids provides an insight into the dynamics of the magma chambers. They are characterized by a fine-grained microgranular texture and are structurally massive with chilled margins towards the MME rims. Darker and finer chilled margins are products of the interaction between felsic host rocks and their mafic magmatic enclaves (e.g., Bussy, 1991; Didier and Barbarin, 1991). MMEs could be restites (Chappell et al., 1987; Chappell and Stephens, 1988; Chen et al., 1998), or they could be inclusions of mantle-derived mafic magma (Eichelberger, 1980; Vernon, 1984; Holden et al., 1987; Bonin, 2004; Yang et al., 2006), remnants of lower crustal mafic rocks (Kepezhinskas et al. 1997; Maas et al. 1997), or products



Fig. 12. Schematic examples of processes that can influence the shape of the CSD (after Higgins, 2006, modified by Vinet & Higgins 2010).

(a) Increase of the cooling rate; (b) decrease in line slope resulting in an increasing in growth and residency time in the magma chamber; (c) different segregation effects or crystal accumulations in the system; (d) effects of the presence of coarse grains on the CSD diagram; (e) mixing of two magmas with different crystal densities or coarse-grained crystals. Each of these processes can have an effect on CSD through changing some characteristics, such as residency time.

of magma mixing (e.g. Didier and Barbarin, 1991; Perugini et al., 2003). MMEs are identified here on the basis of their dark color, grain size and textural properties. K-feldspar and plagioclase megacrysts in the MMEs (Fig. 4), together with the presence of feldspars and quartz that were entrained from the host granitoids to the MMEs, suggest a hybrid system formed by the mixing of two distinct end-member magma compositions (e.g., Barbarin, 1990; Hibbard, 1991; Yang et al., 2015). Additionally, the presence of acicular apatite in the MMEs suggests that the mafic magma was injected into the felsic magma, and crystallization in a quenched environment occurred as a result of the mingling of small volumes of hot mafic melt with cooler granitoid magma (e.g., Vernon, 1984; Chen et al., 2009, 2016). Another notable point for magma mixing can be zoning in plagioclase, which is expressed as An% content, and can be governed by a number of factors, among which are changes in temperature, water fugacity (e.g. Loomis and Welber, 1982; Pietranik and Koepke, 2014), and magma composition associated with recharging by more mafic magma (e.g. Tepley et al., 1999; Ginibre et al., 2002; Browne et al., 2006; Nicotra and Viccaro, 2012).

As already noted, the maximum difference in X_{An} in Tafresh plagioclase is more than 30 mol%. Variations in An content, which can be very large and abrupt, are due to fractionation, and/or the differences in sources, assimilation of wall rocks, and other magmatic processes

(see Yang and Lentz, 2005) and variation in the crystallization condition. Ghiorso et al. (1983) showed that if temperature varies by 10°C, the X_{An} value would change by 1 mol%. Therefore, the sample would require an unreasonably high change in temperature of 300°C if temperature alone had caused the variation in X_{An} . Ghiorso et al. (1983) further showed that if pressure varied by 1 kbar then the An value would change by 1.4 mol%. Longhi et al. (1993) specified that a 1 kbar increase in pressure, acting alone, leads to a 1 mol% decrease in X_{An} . Similarly, the 30 mol% range in X_{An} would correspond to an absurdly large pressure range of about 30 kbar. It is more plausible that a process of magma mixing has been responsible in this instance.

Finally, and as a result, the presence of the sieve texture, microgranular enclaves, oscillatory zoning in plagioclase crystals and the changes in the slope of the CSD line together indicate the possibility of magma mixing in the Tafresh granitoids. Generally, magma mixing has been shown to have a large influence on the mineralogical, textural and geochemical conditions of arc-related rocks (Tepley et al., 1999).

6 Conclusions

The Tafresh granitoids, which are part of the Urumieh-Dokhtar Magmatic Arc in central Iran, are I-type, metaluminous, and calc-alkaline.

The chemical compositions of the amphiboles, which are calc-alkaline, indicate a high oxygen fugacity, which is consistent with a mixed crust-mantle magma source. On the basis of Altot contents in the amphiboles of Tafresh granodiorites, the final crystallization is indicative of a low pressure of crystallization.

Calculated residence times for Tafresh plagioclase crystals is less than 1,000 years, which supports field evidence that these plutons were emplaced at shallow depths. The shape of the CSD curves in the samples are consistent with the standard magma mixing CSD patterns.

The presence of sieve texture, microgranular enclaves (MMEs), oscillatory zoning in plagioclase crystals and variations in the slopes of CSD lines are a consequence of magma mixing in the Tafresh granitoids.

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